

Microsoft Excel - extPpalCS39

Fichier Édition Affichage Insertion Format Outils Données Fenêtre ? Adobe PDF Tapez une question

Répondre en incluant des modifications... Terminer la révision...

J5 fx 1

	A	B	C	D	E	F	G	H	I	J	K	L	M	N	O	P
1	type	HD	mV	MV	Teff	RQ	dir	count	11-16+3k<C<20k;3k<A<7k	PropA1	PropA2	PropA3	PropA4	PropA5	PropA6	
2	sol-like	42807	6.4	5.1	5750	Anticentre		3261		0						
3	sol-like	42618	6.9	5	5800	Anticentre	6031		1							
4	sol-like	46241	5.8	6.35	4900 near SRa02	Anticentre	4612 Too cold			0						
5	sol-like	165401	6.8	4.9	5850	Centre	9227			1						
6	sol-like	175726	6.7	4.6	6100 SRc01	Centre	2286 NO			0						
7	sol-like	168443	6.9	4	5600 2planets	Centre	10400			1						
8	off-MS sol-like	57006	5.9	2	6030	Anticentre	6352			1						
9	Red Giant	43023				Anticentre	3918			0						
10	Red Giant	47220				Anticentre	4614			1					1	
11	Red Giant	40726				Anticentre	1181 NO			0						
12	Red Giant	166284				Centre	8794			1						
13	Red Giant	184013			Very east	Centre	14628 NO			0						
14	Red Giant	184297			Very east	Centre	9173 far east			0						
15	Cluster/Red (Berkeley 32)				In exo-field	Anticentre	8223 NO Bright sta			0						
16	Cluster/Red (NGC2296)				In exo-field	Anticentre	7200			0						
17	Cluster/Red (NGC6633)				In seismo-fiel	Centre	9088			1						
18	Cluster/Red (NGC6705)				In exo-field	Centre	24409 NO			0						
19	O-stars	52266			near LRa02	Anticentre	7451			0						
20	O-stars	46573				Anticentre	4861			1					1	
21	O-stars	47129			SRa02	Anticentre	6289			1					1	
22	O-stars	46150			SRa02	Anticentre	4832			0						
23	O-stars	166802				Centre	9348			0						
24	O-stars	172275			IV=9.35	Centre	29828 NO			0						
25	O-stars	BD-08 4617			IV~9.3	Centre	9021			0						
26	O-stars	173783			IV~9.4	Centre	25213 NO			0						
27	B-sGiant	47240			B1 I	Anticentre	4294			1					1	
28	B-sGiant	46769			B8I	Anticentre	6054									

Feuil1 Feuil2 Feuil3

Prêt

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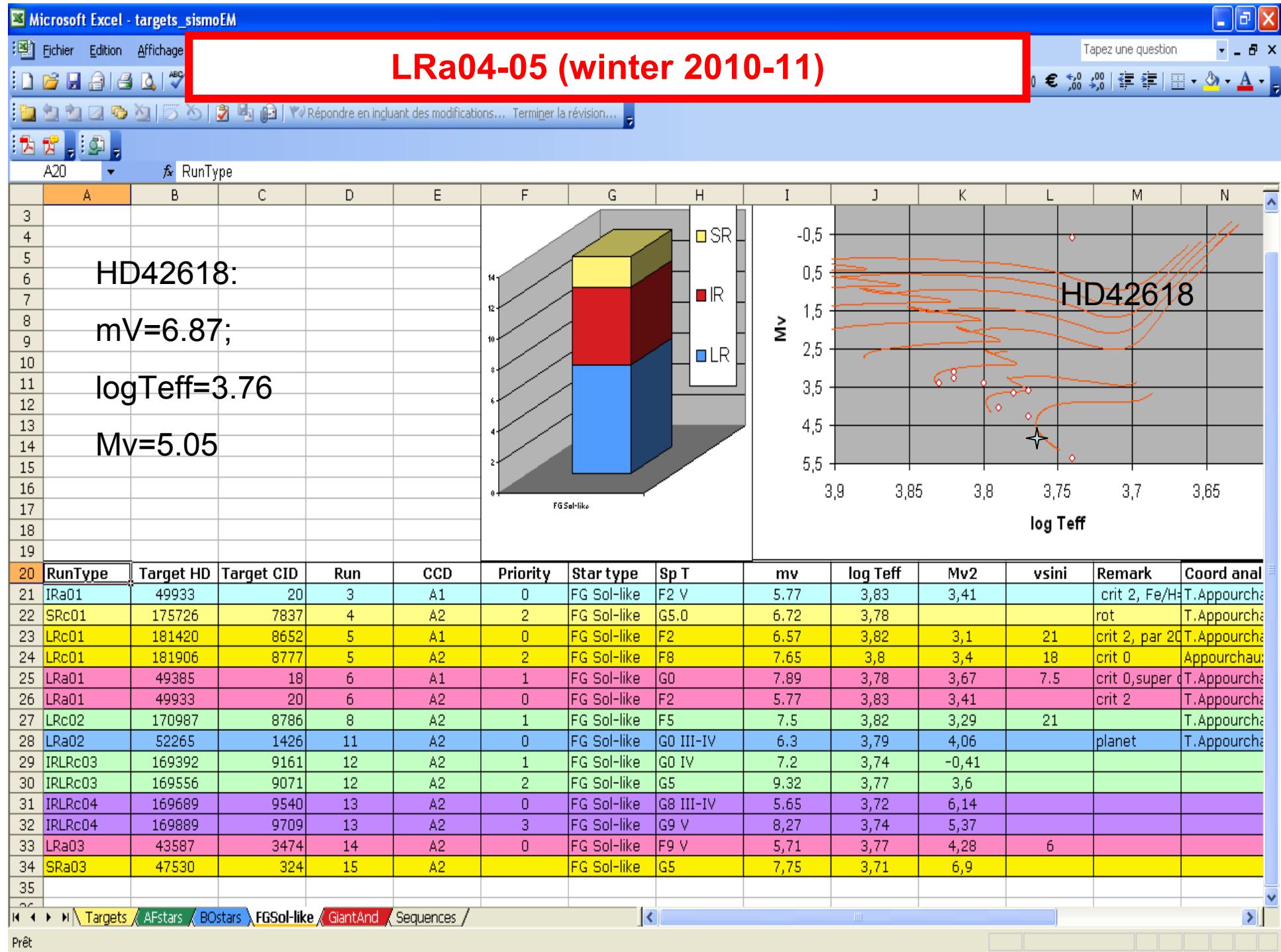
Répondre en incluant des modifications... Terminer la révision...

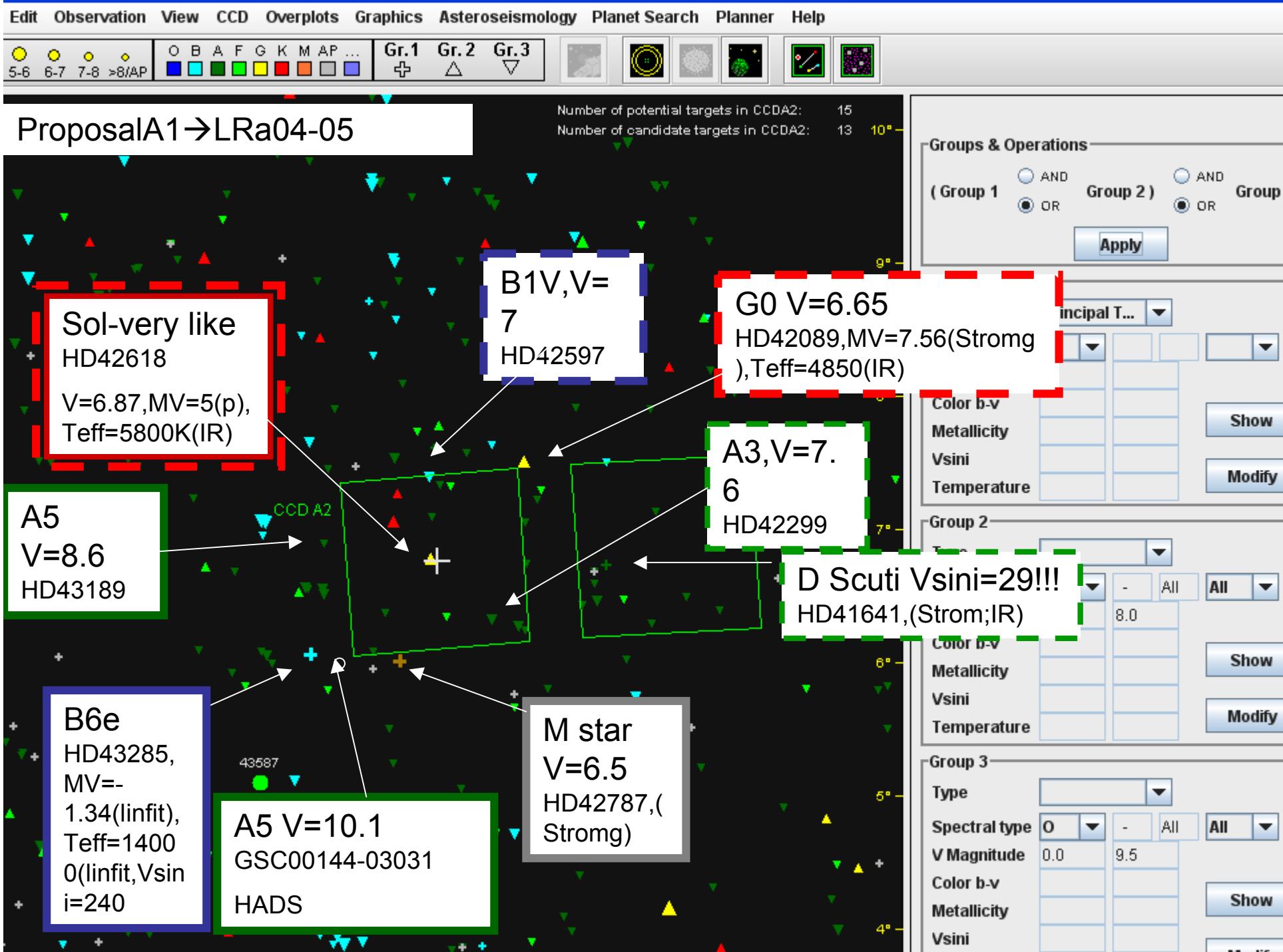
J5 f 1

	A	B	C	D	E	F	G	H	I	J	K	L	M	N	O
14	Red Giant	184297				Ivery east	Centre	9173	far east	0					
15	Cluster/Red	CBerkeley 32				in exo-field	Anticentre	8223	NO Bright sta	0					
16	Cluster/Red	CNGC2236				in exo-field	Anticentre	7200		0					
17	Cluster/Red	CNGC6633				in seismo-fiel	Centre	9088		1					
18	Cluster/Red	CNGC6705				in exo-field	Centre	24409	NO	0					
19	O-stars	52266				near LRa02	Anticentre	7451		0					
20	O-stars	46573					Anticentre	4861		1				1	
21	O-stars	47129				SRa02	Anticentre	6289		1			1		
22	O-stars	46150				SRa02	Anticentre	4832		0					
23	O-stars	166802					Centre	9348		0					
24	O-stars	172275				IV=9.35	Centre	29828	NO	0					
25	O-stars	BD-08 4617				IV~9.3	Centre	9021		0					
26	O-stars	173783				IV~9.4	Centre	25213	NO	0					
27	B-sGiant	47240				B1 I	Anticentre	4294		1			1		
28	B-sGiant	46769				B8I	Anticentre	6054							
29	Beta Cep	48553				prio 2	Anticentre	5815		0					
30	Beta Cep	171305				prio 1	Centre	15802	NO	1					
31	HgMn	46886				prio 1	Anticentre	4459		0					
32	HgMn	45975				prio 2	Anticentre	5672		0					
33	HgMn	47278				prio 4	Anticentre	4254		1					
34	HgMn	55362				prio 5	Anticentre	8851		0					
35	HgMn	173673				prio 3	Centre	25799	NO	0					
36	Be stars	49330				LRa01	Anticentre			0				0	
37	Star+planet	46375				SRa02 +plan	Anticentre			1			1		
38	Cluster	NGC2264					Anticentre	3889		1					
39	Cluster	NGC2244					Anticentre	4132	to many bright	1					
40	EXO	CoRoT-7b				LRa01	Anticentre	8049		1				1	
41	EXO	CoRoT-9b				SR...	Centre	6985		1					
42	type	HD	mV	MV	Teff	RQ	dir	count			PropA1	PropA2	PropA3	PropA4	PropA5

Feuil1 \ Feuil2 \ Feuil3 \

Prêt





Then,  
in the Anticentre direction

# Thierry Morel (G. Alecian)

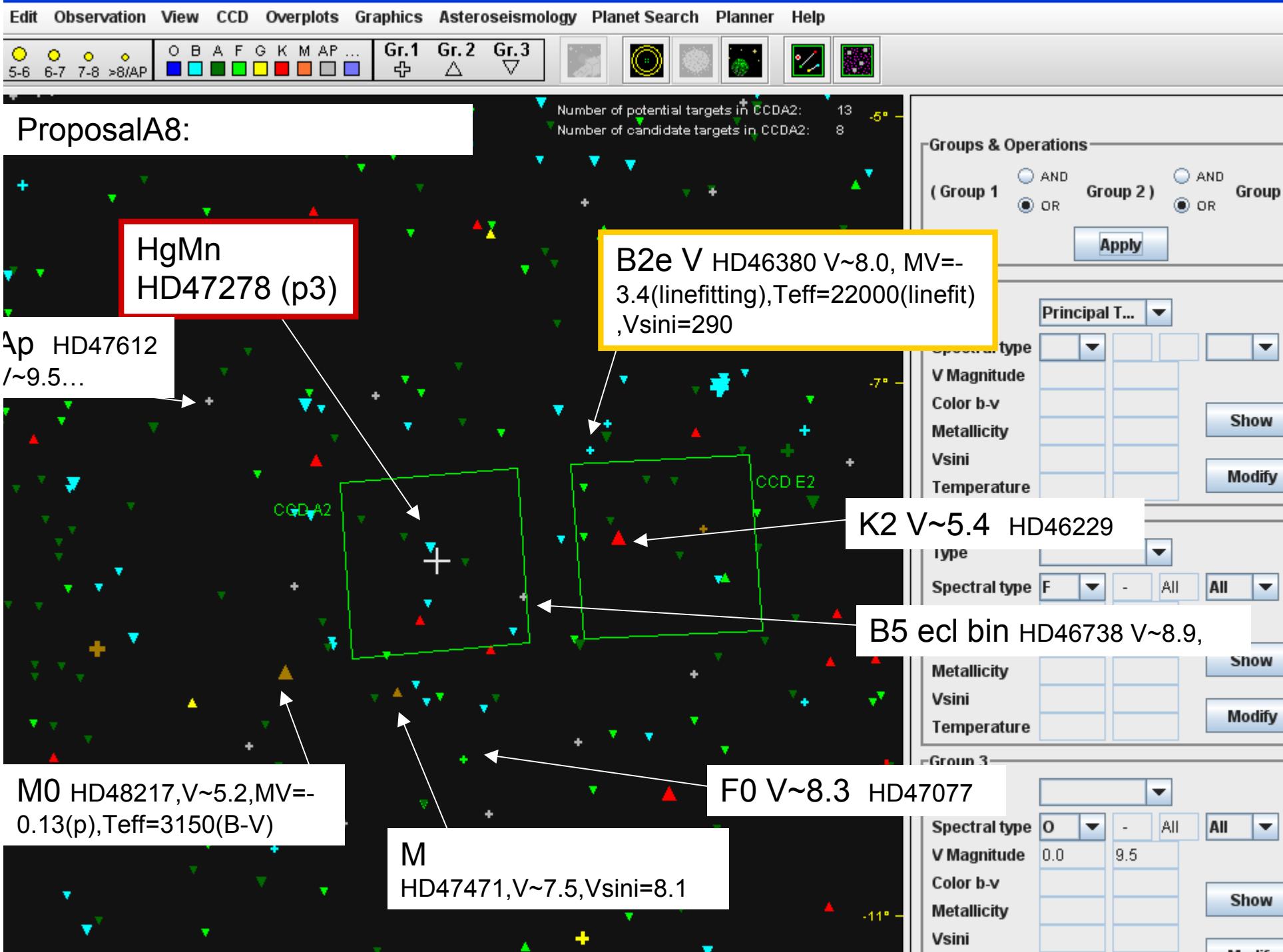
# HgMn stars

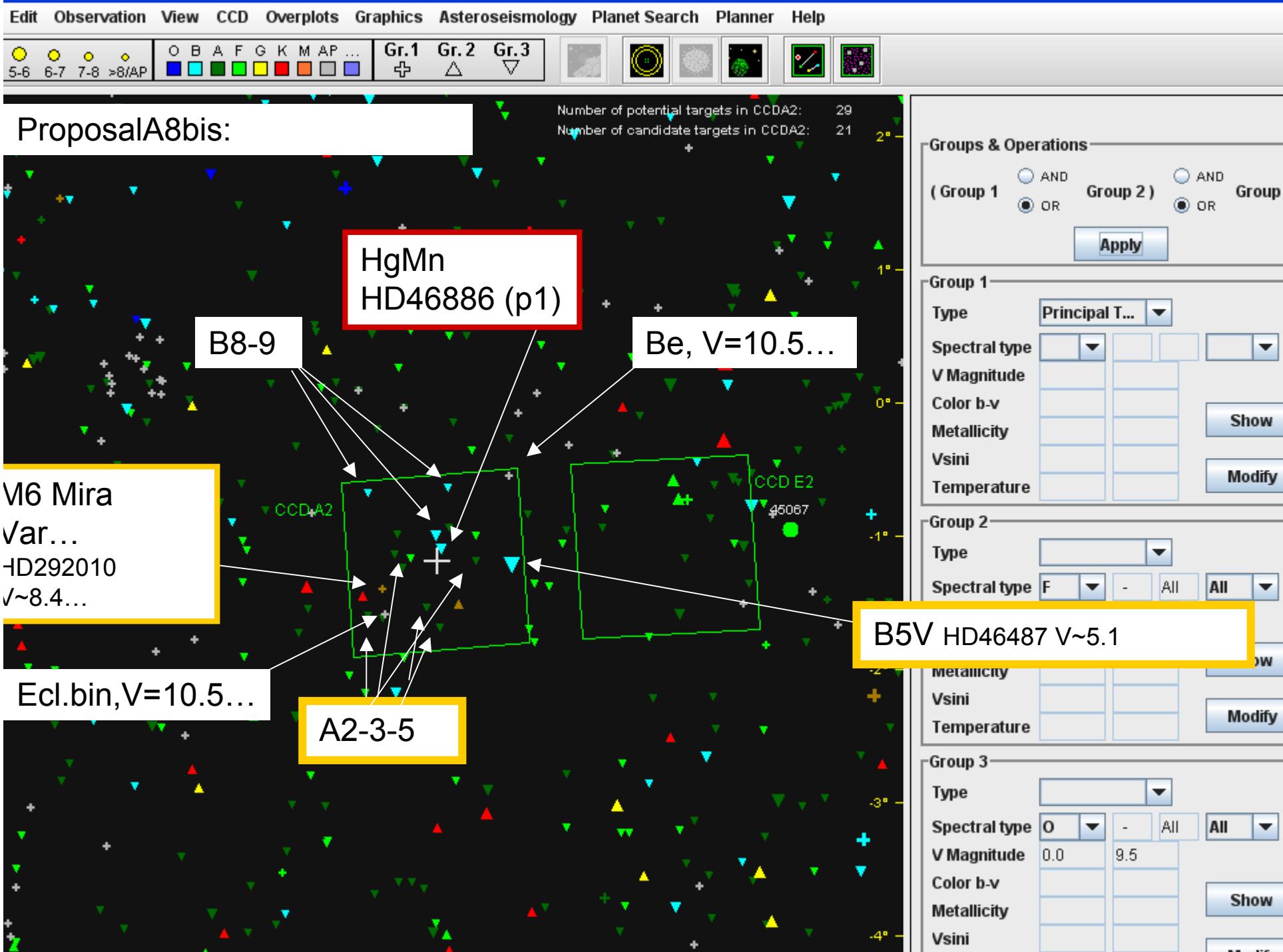
Table 1: List of targets.

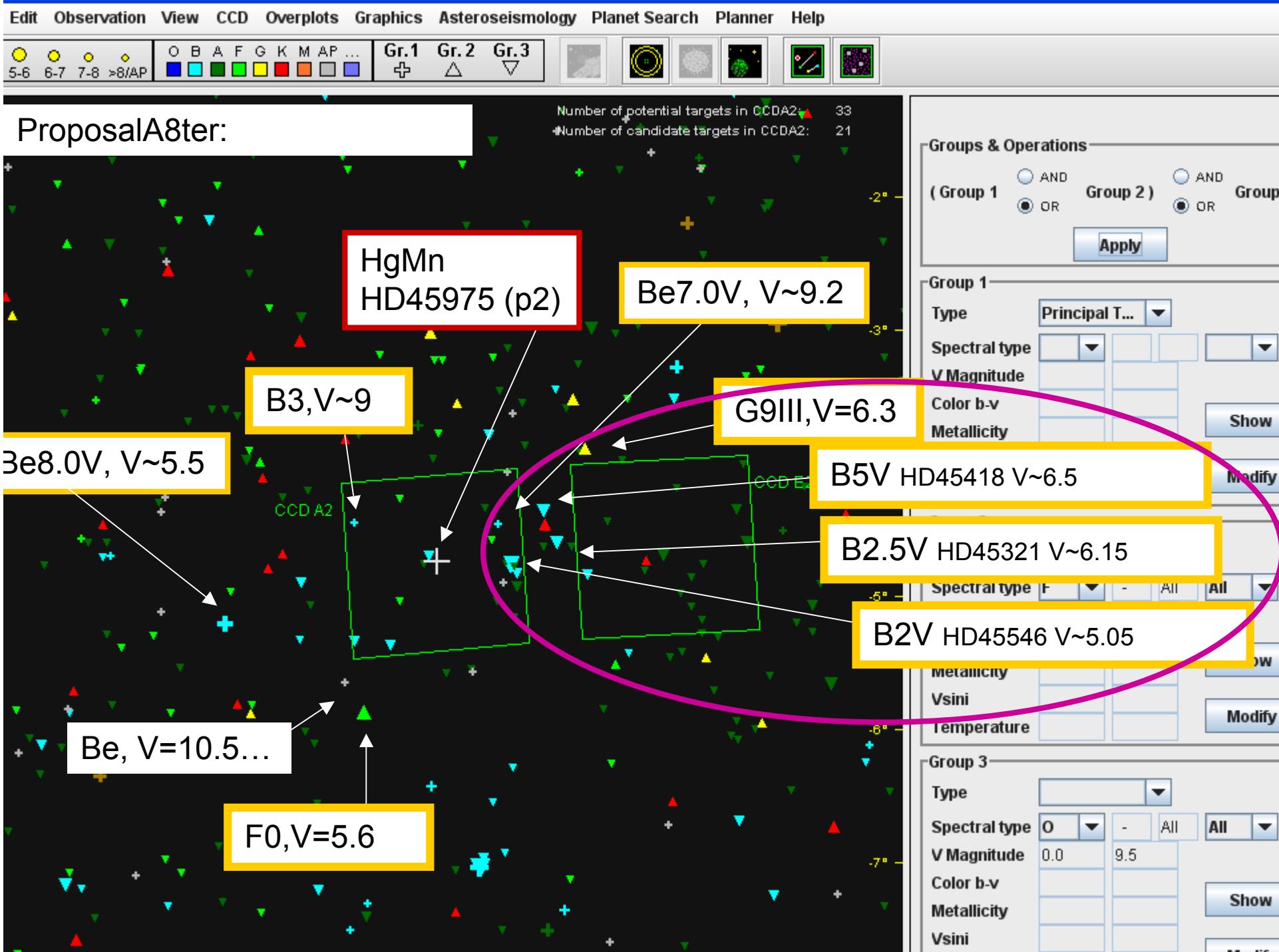
Priority	Target	Sp.T.	$\alpha$ (J2000)	$\delta$ (J2000)	$V$ (mag)	$v \sin i$ (km s $^{-1}$ )	$T_{\text{eff}}$ (K), $\log g$
highest	HD 46886	B9	06:35:44.8	-01:05:48	7.95	21	12 900, 3.8
	HD 45975	B9	06:30:27.5	-04:41:49	7.46	62	12 500, 4.0
lowest	HD 47278	B9	06:37:32.0	-08:14:08	7.23	37	11 500, 4.1

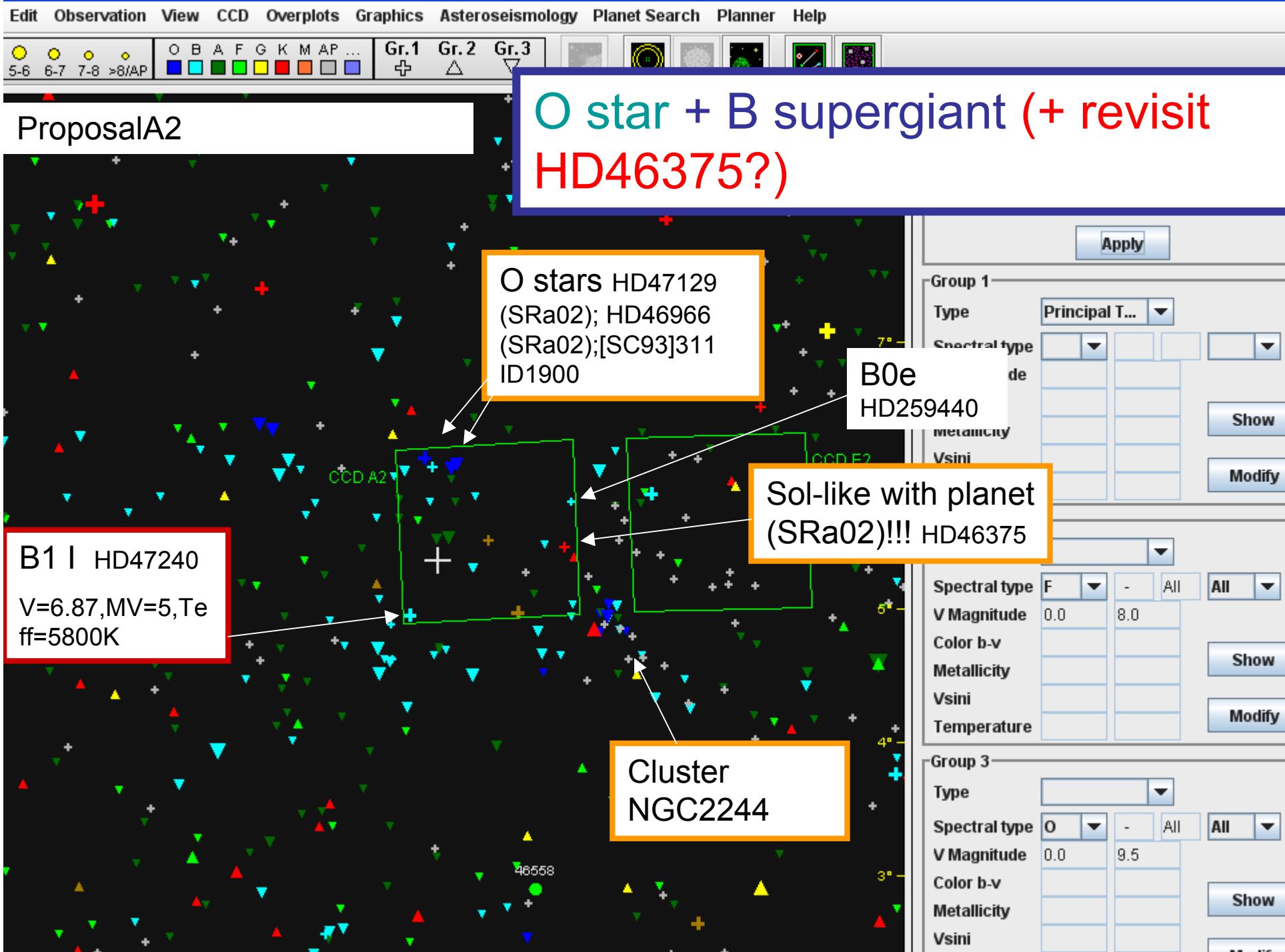
Table 2: Mean abundances (on the scale in which  $\log \epsilon[\text{H}] = 12$ ) and line-to-line scatter. The number of spectral lines used is given in brackets. The first column gives the solar abundances of Asplund et al. (2009, ARA&A, 47, 481).

	Sun	HD 45975	HD 46616	HD 49886	HD 47278	HD 53004	HD 55362	HD 46886
$T_{\text{eff}}$ (K)		12500	16000	13000	11500	11600	13000	12900
$\log g$		4.0	4.0	4.0	4.1	4.0	4.0	3.8
$\xi$ (km s $^{-1}$ )		2	0	0	2	2	2	0
$v \sin i$ (km s $^{-1}$ )		$62 \pm 2$ (18)	$7.7 \pm 1.7$ (38)	$8.50 \pm 0.8$ (84)	$37 \pm 2$ (23)	$51 \pm 8$ (20)	$54 \pm 5$ (16)	$21 \pm 2$ (48)
$\log \epsilon(\text{C})$	8.43	7.60 (1)	$7.98 \pm 0.20$ (3)	$6.97 \pm 0.80$ (2)			7.82 (2)	$8.15 \pm 0.01$ (2)
$\log \epsilon(\text{N})$	7.83		$8.29 \pm 0.20$ (5)					8.37 (1)
$\log \epsilon(\text{O})$	8.69		$9.14 \pm 0.30$ (2)					
$\log \epsilon(\text{Mg})$	7.60	$7.45 \pm 0.12$ (2)	$7.37 \pm 0.17$ (2)	$6.60 \pm 0.11$ (4)	$7.22 \pm 0.12$ (2)	$7.42 \pm 0.06$ (2)		$7.13 \pm 0.22$ (2)
$\log \epsilon(\text{Al})$	6.45							8.12 (1)
$\log \epsilon(\text{Si})$	7.51	$7.00 \pm 0.38$ (2)	$8.21 \pm 0.16$ (3)	$7.49 \pm 0.09$ (3)		7.55 (1)		$7.13 \pm 0.04$ (2)

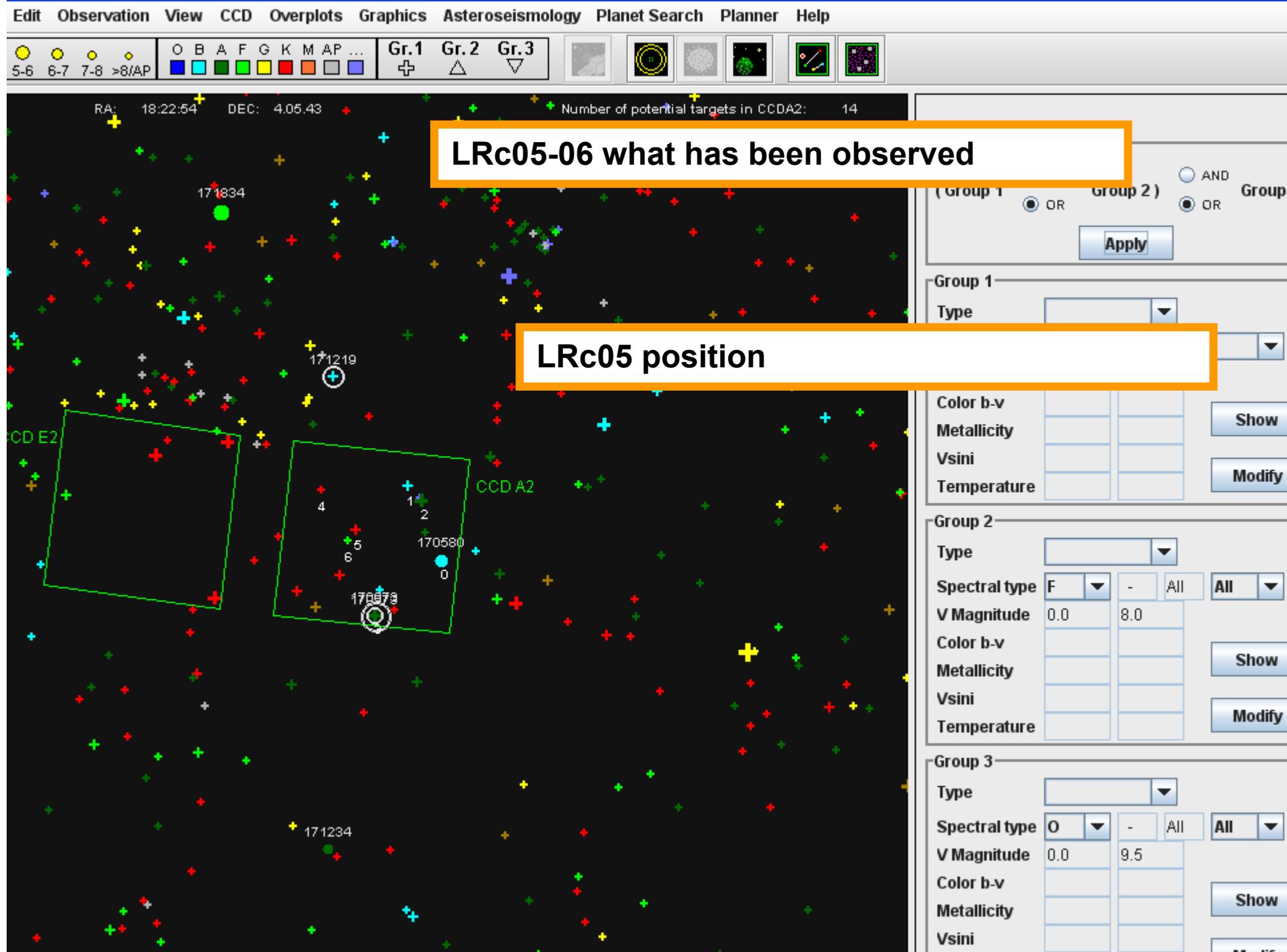


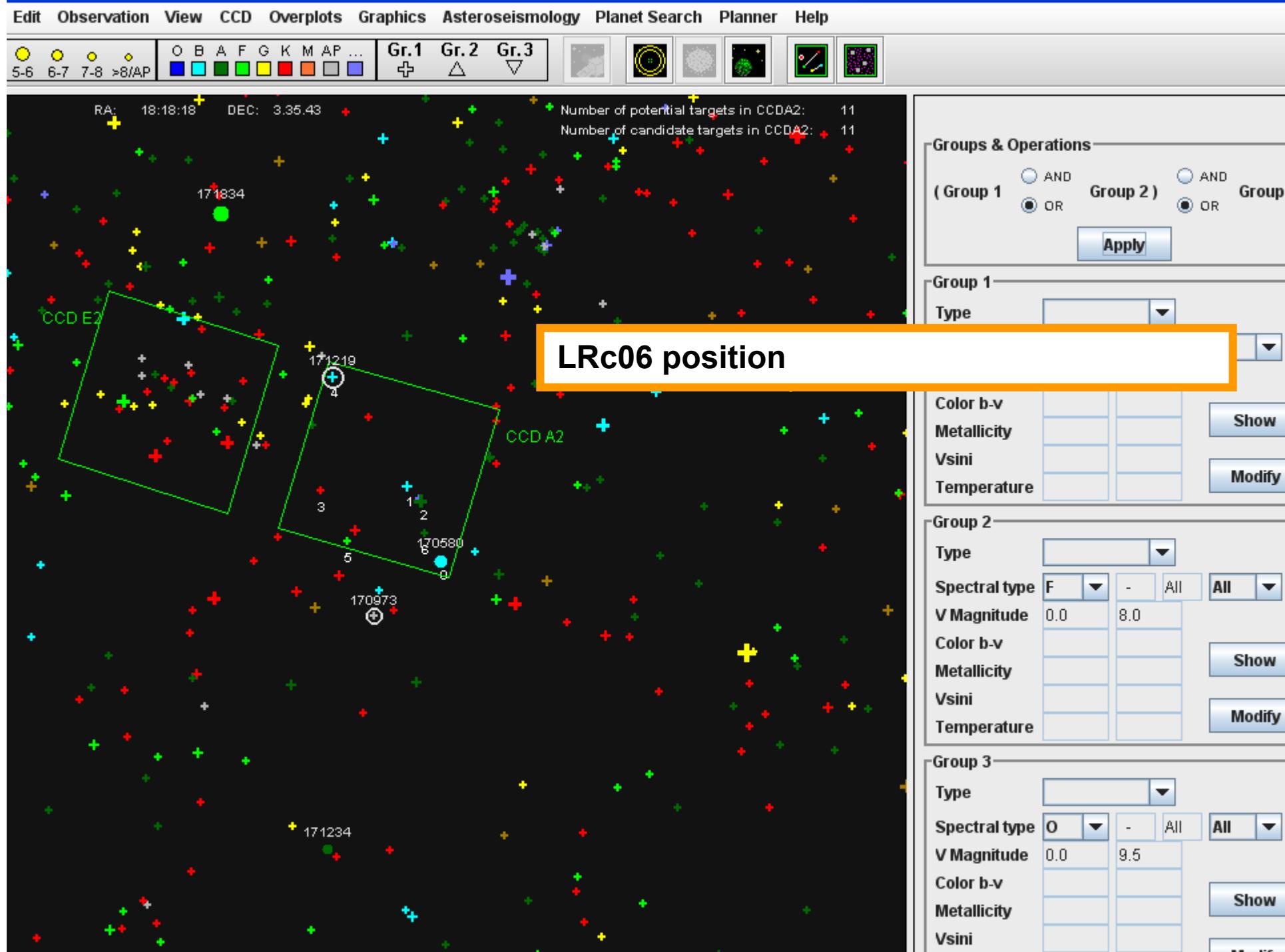












HD171219:p1 B5IIIe, mid-Be star, quite rare in Corot's field of view, will probably show more pulsation modes than cooler Be. In addition, strong double-peaked emission at Halpha and Hbeta -> information on disk configuration and star inclination (constrain the modelling)



HD171264.K0III giant:p3 low L, which would imply a nu\_max well over 250 microHz, which is unfortunately out of reach for CoRoT. At these frequencies, the amplitudes of the peaks are too small to be detected by CoRoT.'

HD 170783: p1 'poorly known, ... cooler than B5... but clearly hotter than the more boring B8/9 stars so still worth the effort...fast rotator...that may complicate matters. interesting object to bridge the slowly rotating SPBs and the pulsating Be stars.'

HD 170699 p2 Delta Sct...fast rotator...but bright (V=6.9), possibility ground-based photometry/spectroscopy

HD 170580:p1 "only star in the hybrid part of the Beta Cep and SPB strip, and so a potential candidate for p- and g-modes simultaneously, ...Moreover very low vsini "

HD 170973:p1 CP star with SrCrSi overabundance ...known rotational period of 18.2 days...possibility for parallel ground based spectroscopy with the goal of Doppler Imaging. ...rather hot (11600K), hence, I have no hope that we might discover rapid pulsations - but who knows...'

Claire M.

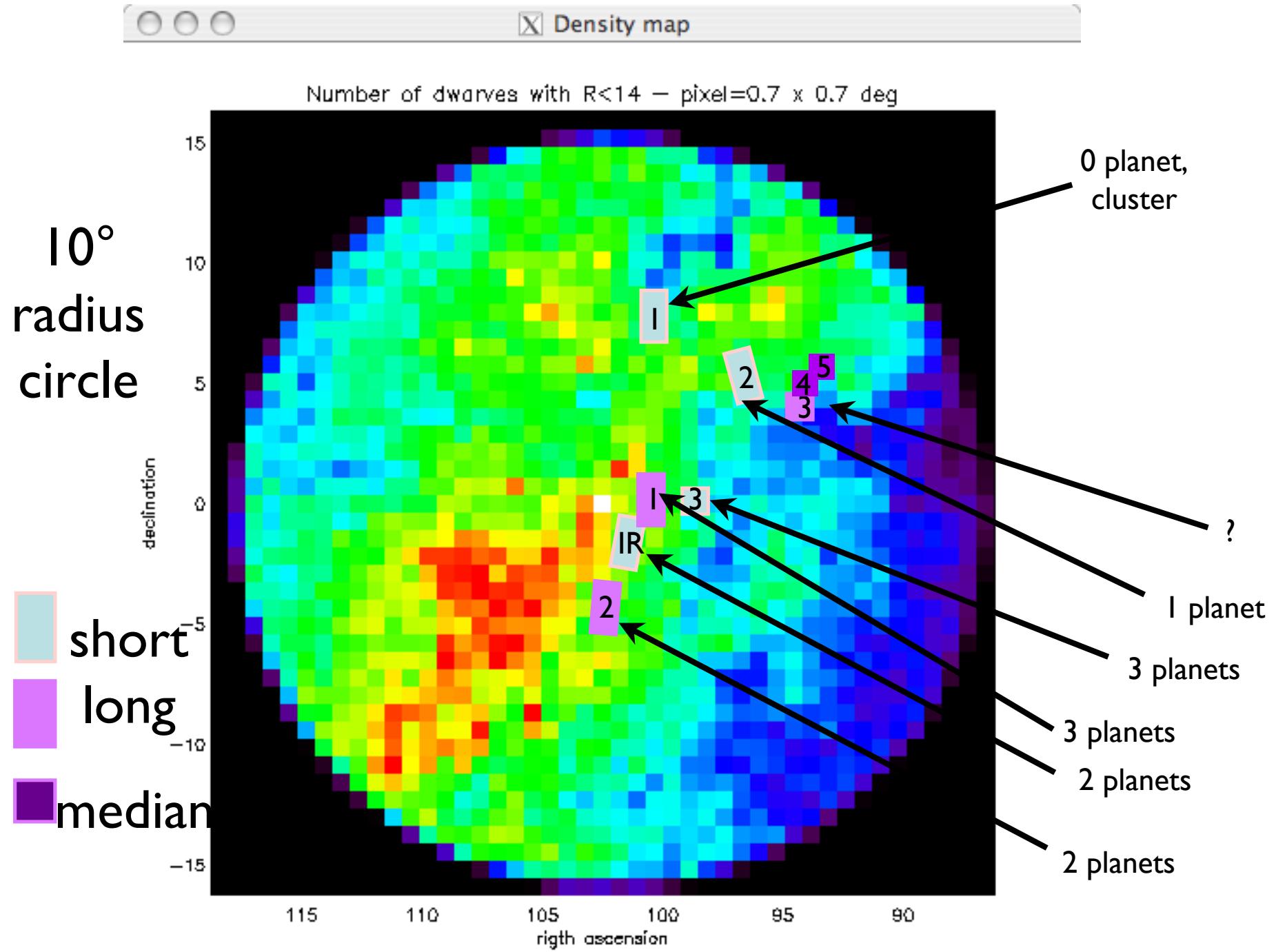
## anticenter fields

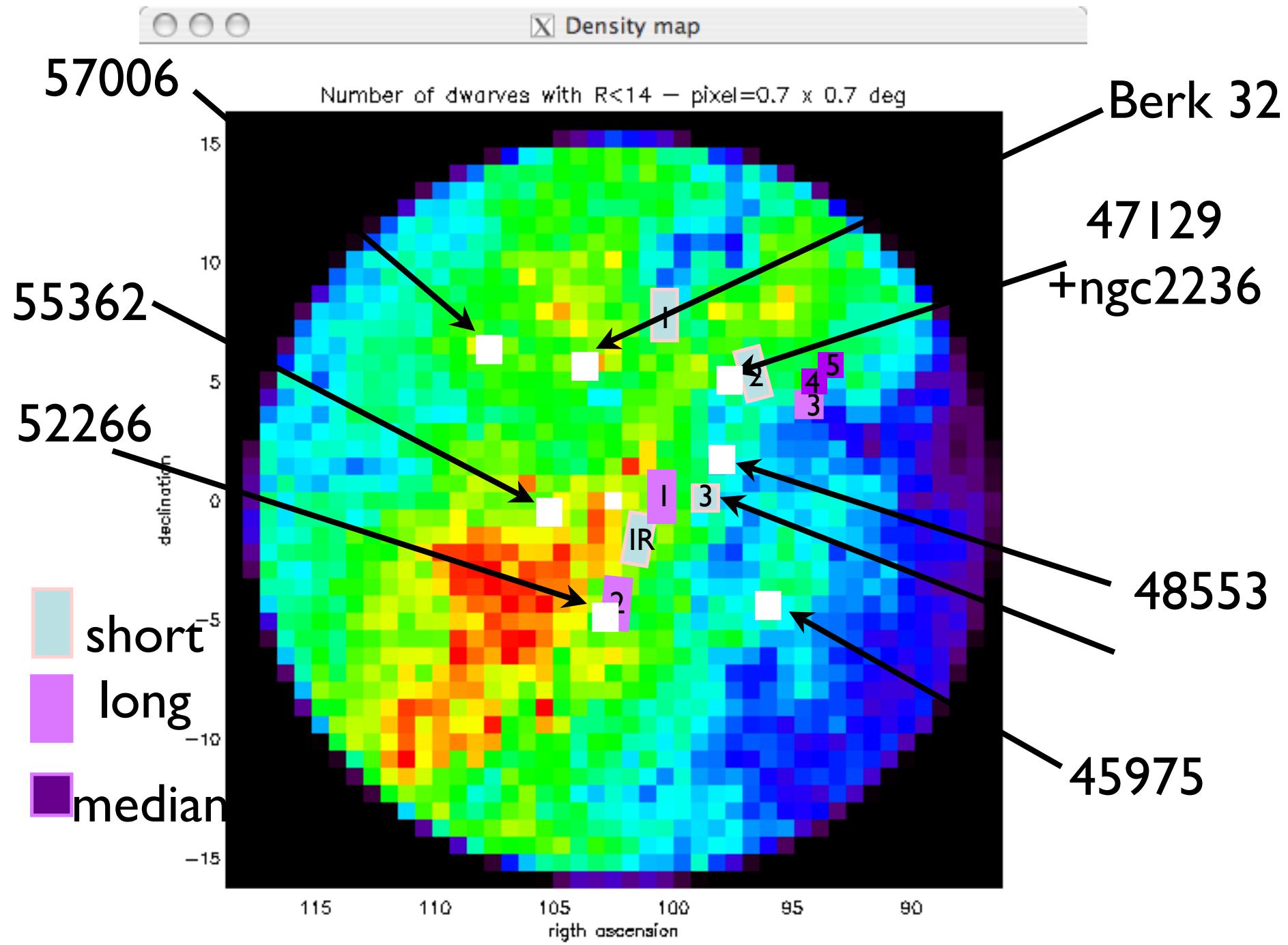
field name	dw r<16	nb planets
SRa03	4270	3!
SRa01	5750	0
SRa02	5550	1
LRa03	6300	0!
LRa02	6380	2
IRa01	7400	2
LRa01	8200	4
LRa04/5	6000	-

(per ccd)

meta name	dw r<16
40726	1181
42807	3261
46241	4612
43023	3918
47220	4614
46150	4832
46573	4861
47278	4254
46886	4459
47240	4294
47278	4254
45975	5678 close to SRa03! (good region)
48553	6289 apparently)
47129	6352
57006	7200
NGC 2236	7451 overlaps LRa02
52266	8049
corot7b	8223
Berkeley 32	8851
55362	contamination?

not enough stars!





stellar density:

avoid yellow, red, dark blue regions in maps

non-covered «green/light blue» regions are all of  
potential interest

preference for 48553, 57006, Berk 32 among sismo fields

Corot7:

a HARPS proposal for reobserving corot7 has been  
put in (results in July); if successful, simultaneous  
corot observations would be a strong plus