



# HIGHLIGHTS OF BRAZILIAN CONTRIBUTIONS TO CoRoT

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CoRoT 44th SC





➢ Brazil Space Agency - CNES

agreement for CoRoT in 2003

■ Brazilian participation in CoRoT:

☀ -- Software

 -- Alcantara ground station

❖ -- Science teams

Thanks to: CNPq, CAPES, INPE, AEB,  
Universities...





## Software engineering

- ◆ Characterization of CoRoT instrument  
(L.P. da Silva)
- ◆ System validation, PSF (V. Parro)
- ◆ PSF + jitter correction (F. Fialho, V. Parro)





# SOFTWARE

## ◆ Characterization of CoRoT instrument (L.P. da Silva)

- 1. **Electronic Bias Level** (function of Temperature)
- 2. **Electronic Gain** (function of aging)
- 3. **PSF** (as a function of CCD position)
- 4. **Spontaneous Generation of e<sup>-</sup>** (function of Temperature)
- 5. **Pixel Response: Non-Uniformity**
- 6. **p<sup>+</sup> Impacts** (as a function of Geo Position)
- ...etc...

(da Silva et al. 2006; Auvergne et al. 2006)

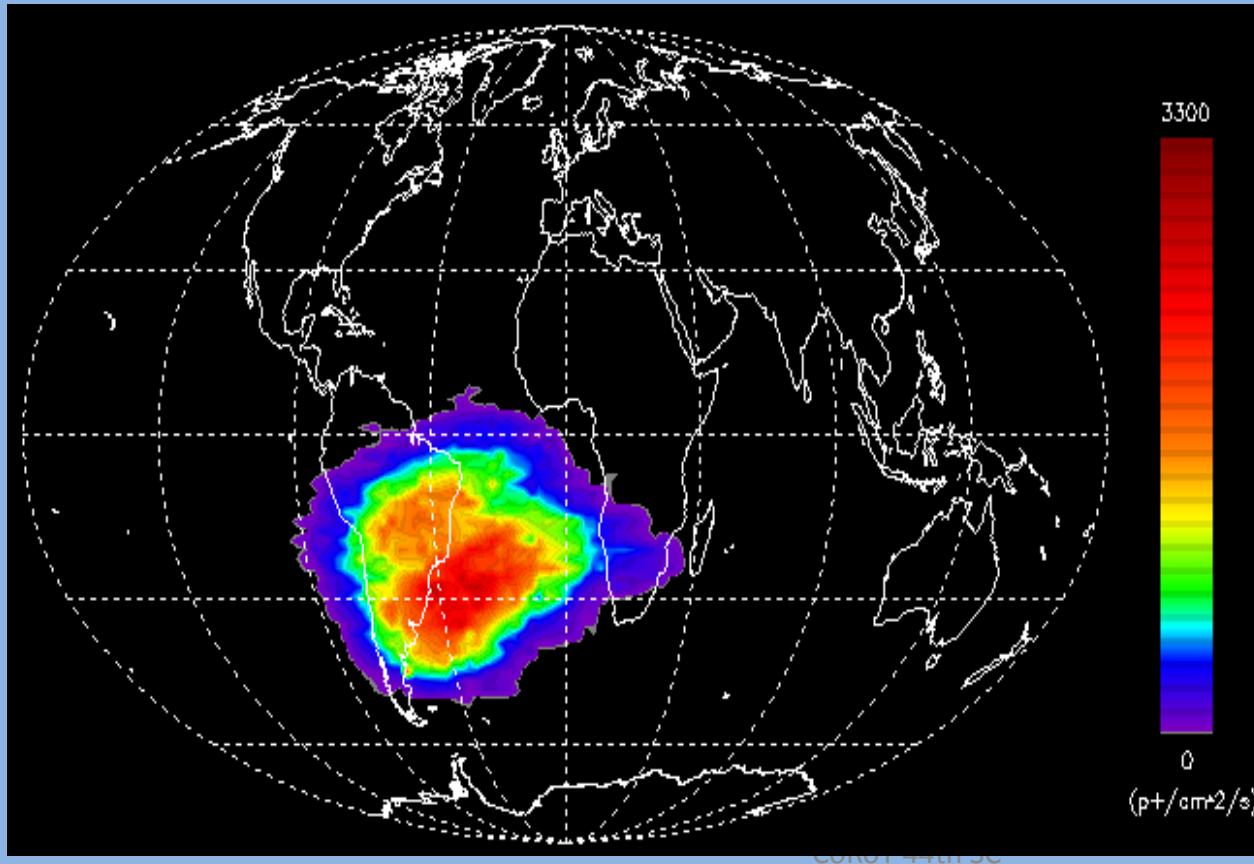
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# SOFTWARE

$p^+$  flux in the focal plane (“orbitography”)



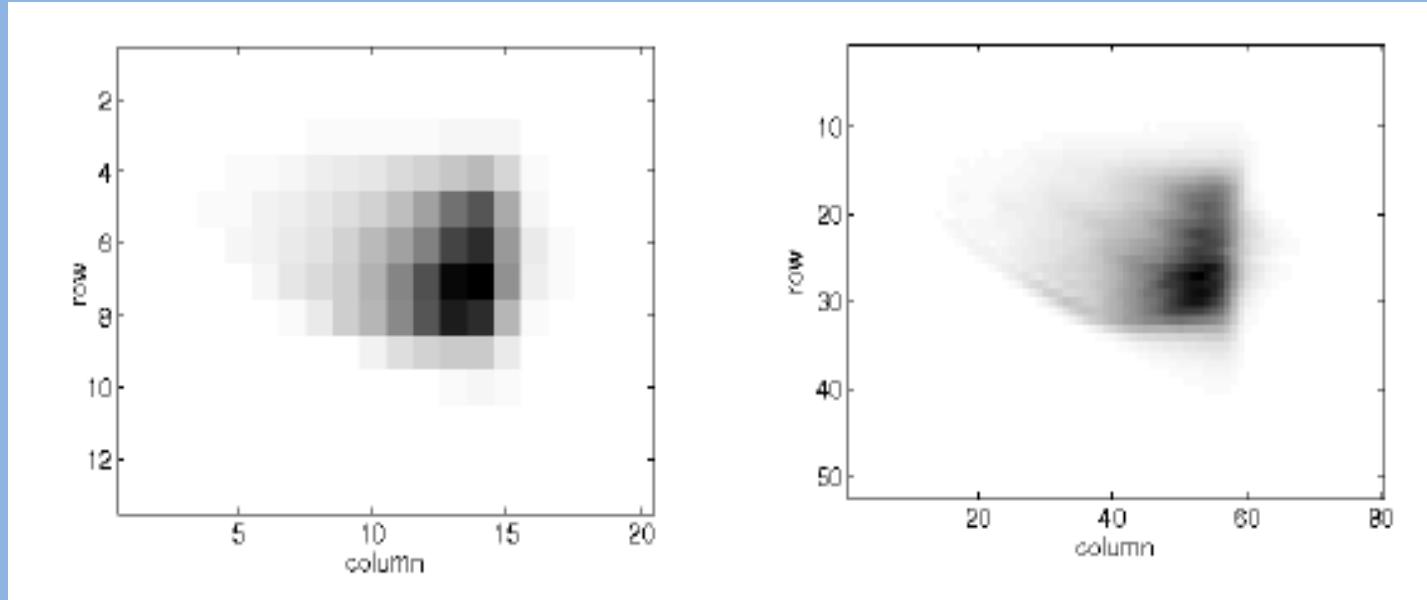
Mapping of the  
**South Atlantic  
Anomaly**  
**(da Silva et al. 2008)**





# SOFTWARE

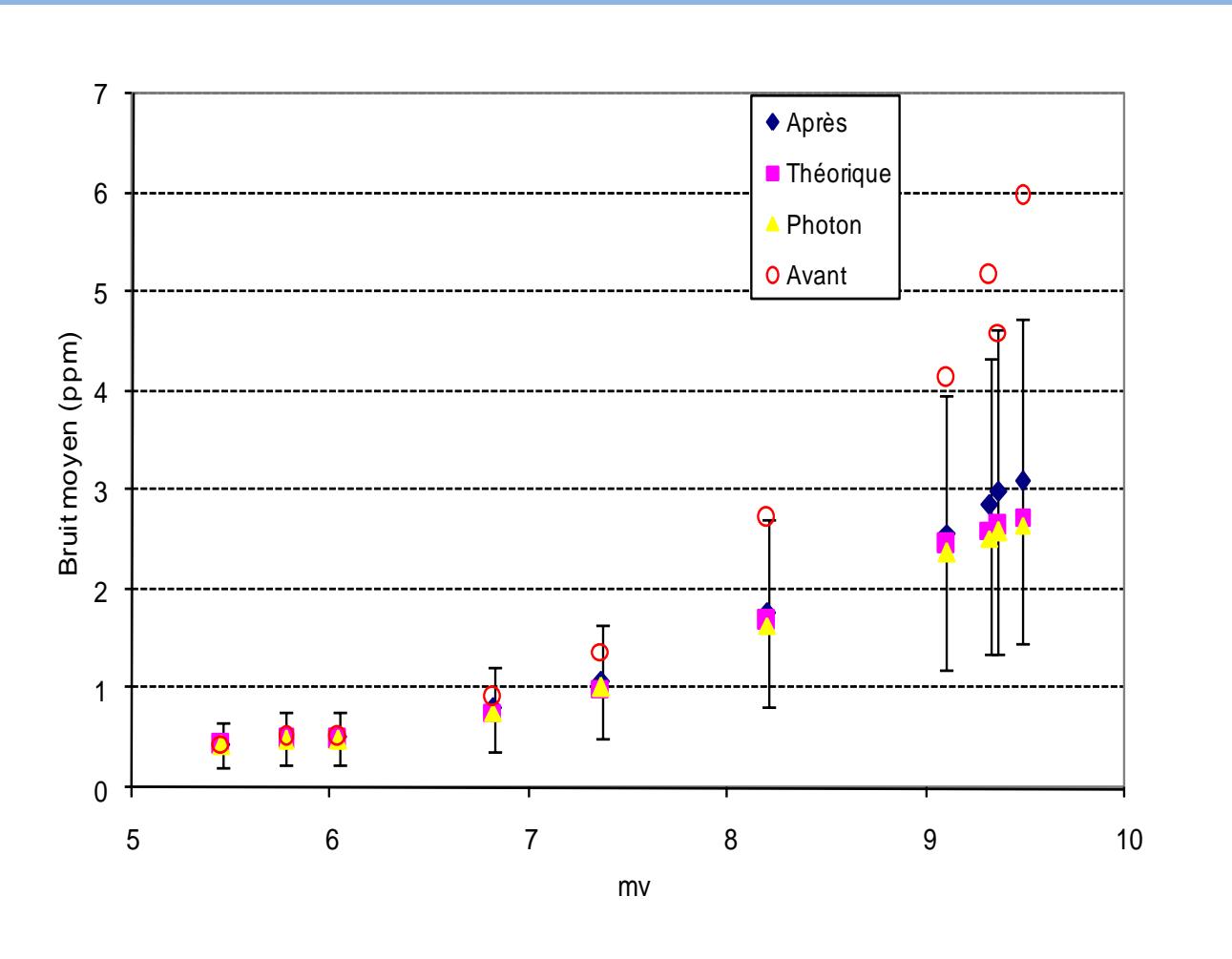
Signal processing and instrument modeling  
for ground segment (PSF +Jitter; Parro &  
Fialho) (Parro et al. 2010)





# SOFTWARE

## Noise correction



CoRoT team

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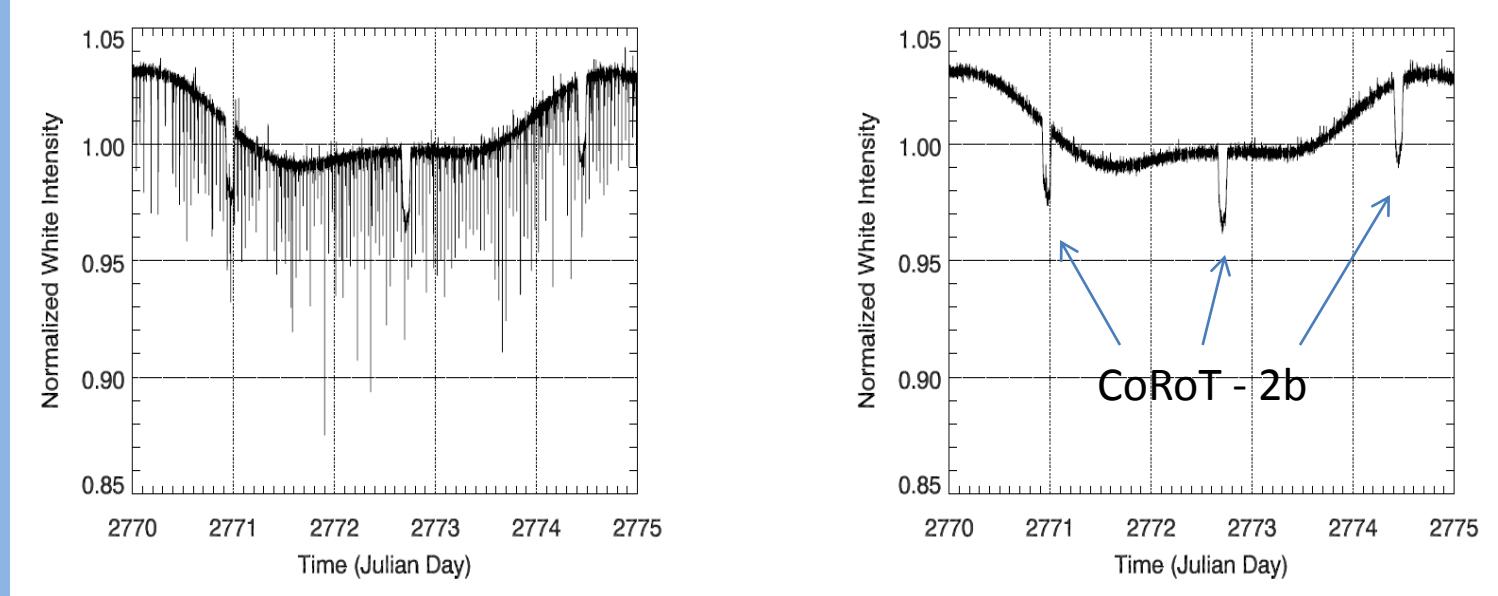




# SOFTWARE

Signal processing and instrument modeling  
for ground segment (PSF +Jitter; Parro &  
Fialho) (Fialho et al. 2009)

Jitter correction for CoRoT-2a light curve.



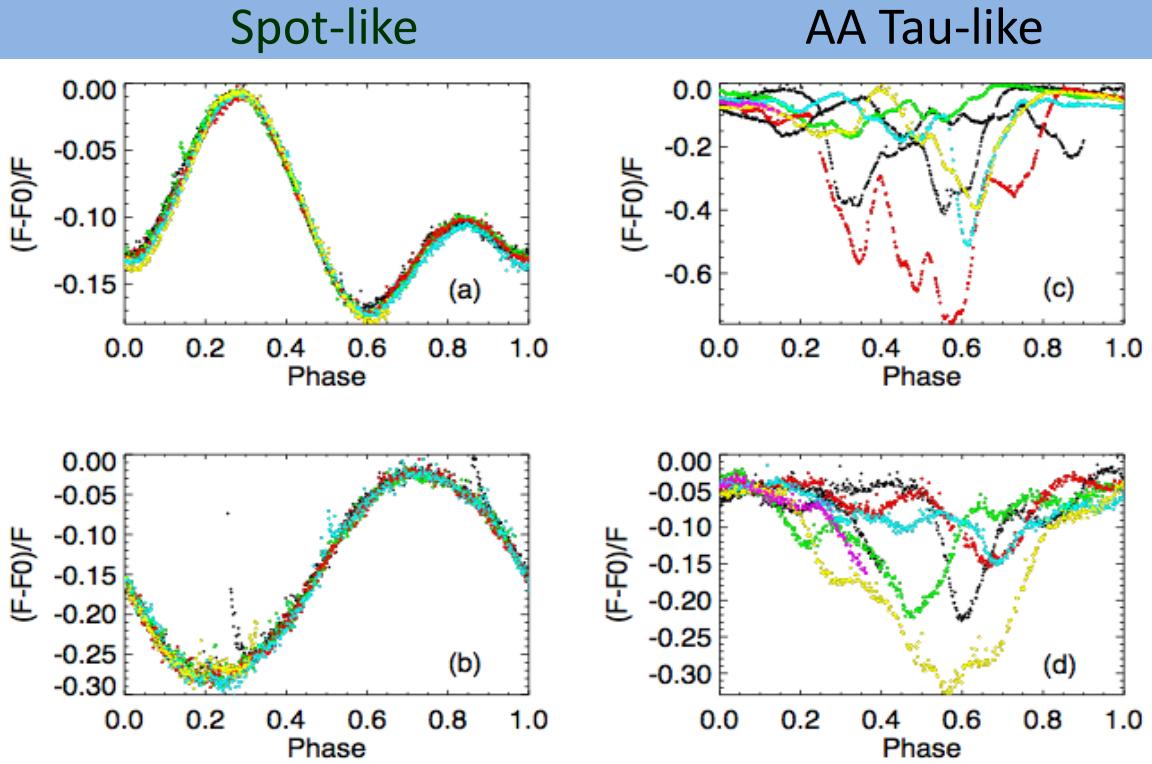


## Science participation

- ☛ Pre-launch observations (La Silla + Brazil)
- ★ Young stars in NGC 2264 (Alencar et al.)
- ★ Dynamics of exoplanets (Ferraz-Mello et al.)
- ★ Stellar seismology (Janot-Pacheco et al.)
- ★ Stellar rotation (Medeiros et. al.)
- ★ Physics of spots (Silva-Valio et al.)
- ★ ...



## Young stars in NGC 2264 (star-disk interaction: CTTSs) (Alencar et al.)



Bouvier et al. 2007;  
Alencar et al. 2012

**Fig. 2.** The four periodic LCs from Fig. 1 folded in phase. Different colors correspond to different cycles.

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# SCIENCE

## Young stars in NGC 2264 (star-disk interaction) (Alencar et al.) (cont)

- Is an inner disk warp a common feature in CTTSs ? Yes, 28% of the observed CTTSs exhibit AA Tau-like light curves.
- Is the dynamical star-disk interaction with a timescale of few rotational periods common among CTTSs ? Yes!
- Is there really dust in the inner disks of the AA Tau-like systems ? Yes, all the AA Tau-like systems all have dust in their inner disk ( $\alpha_{\text{IRAC}} > -2.86$ ).





# SCIENCE

## Dynamics of exoplanets (Ferraz-Mello et al.)

- Co-I  CEST
- Orbit & mass determination from analysis of radial velocity (CoRoT 7b & c: improvement + CoRoT 14b, 16b, 20b, 22b, 23b, 24b & c)
- Tides: determination of shape of planets & of tidal locking.
- analysis of transits time and duration

(Ferraz-Mello 2011; 2011; 2011; 2012; 2012)

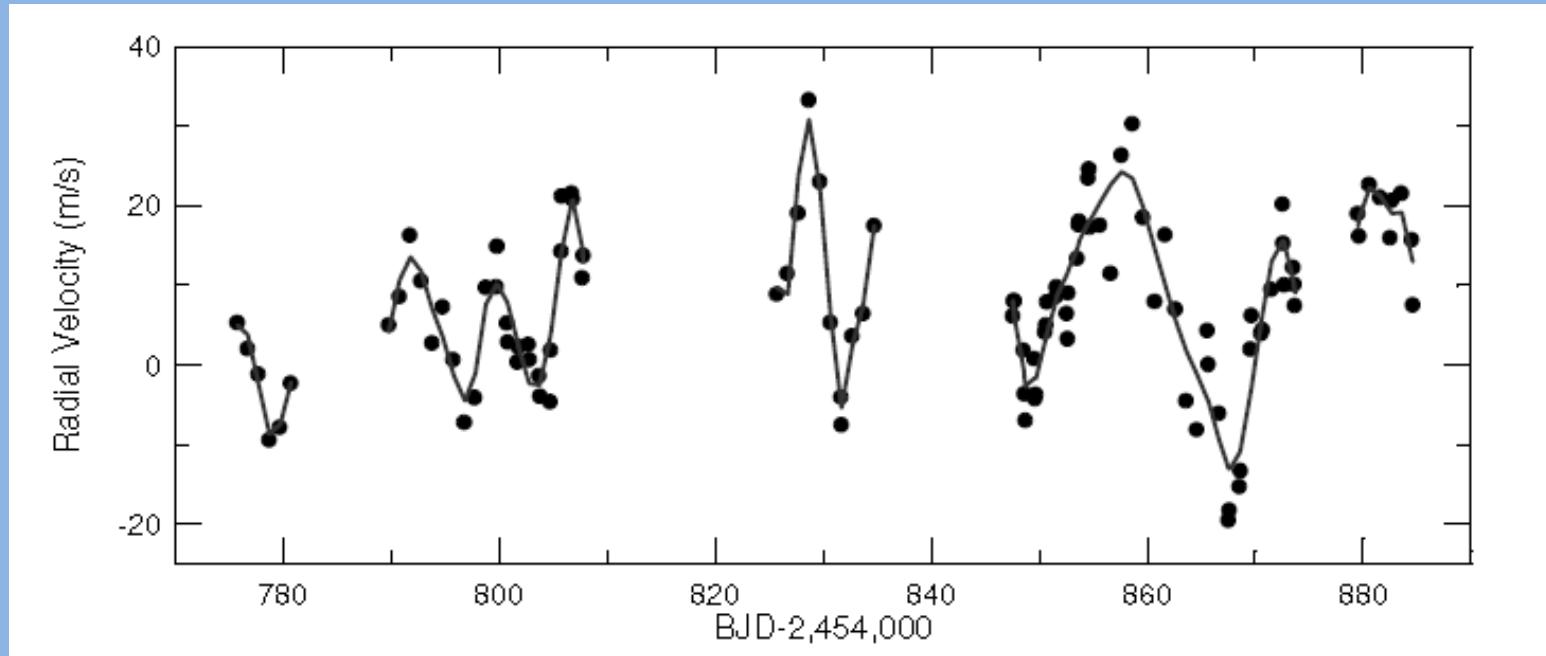




# SCIENCE

## Dynamics of exoplanets (Ferraz-Mello et al.)

**CoRoT 7b** radial velocity curve (HARPS) (Queloz et al. 2009)



( $\sim$  100 nights; orbital period:  $\sim$ 20h)

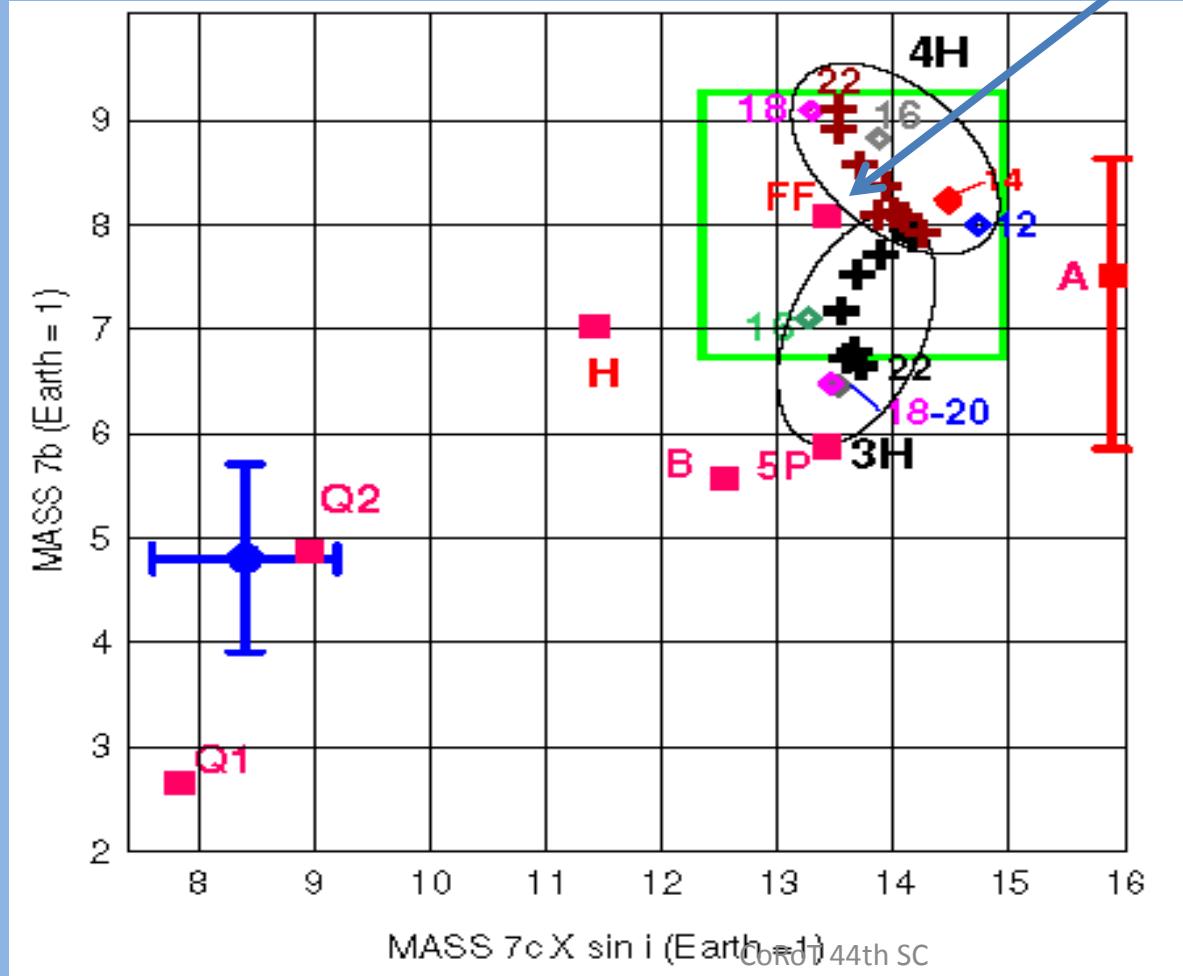




# SCIENCE

## Dynamics of exoplanets (Ferraz-Mello et al.)

CoRoT 7b&c: all results put together





# SCIENCE

## Stellar seismology (Janot-Pacheco et al.)

Pulsation analysis of hot stars (Be,  $\beta$  Cep) and solar-like stars ( $M > 1.3 M_{\text{Sun}}$ ) + comparison with CESAM & FILOU codes →→

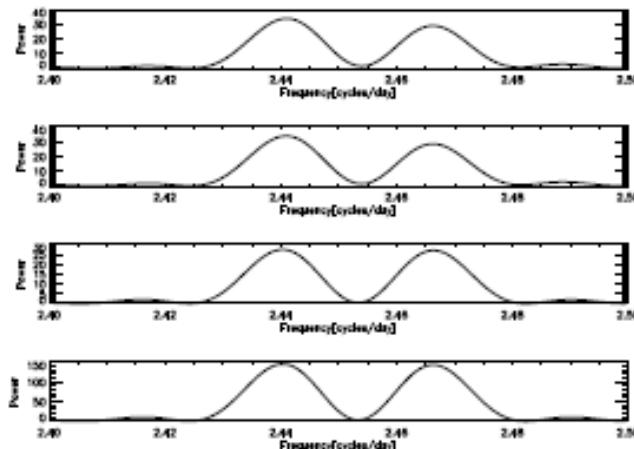
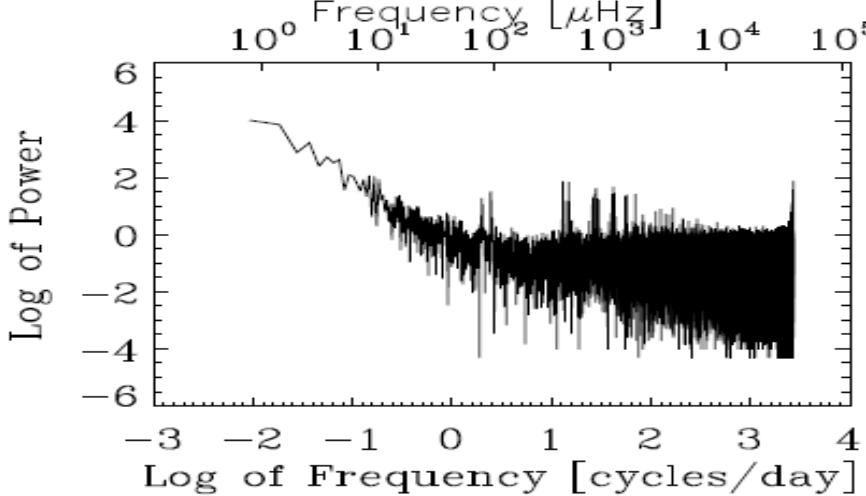
→→ Diagnosis of **stellar structure** and internal rotation profile

(Andrade et al. 2012; Janot-Pacheco et al. 2012 )



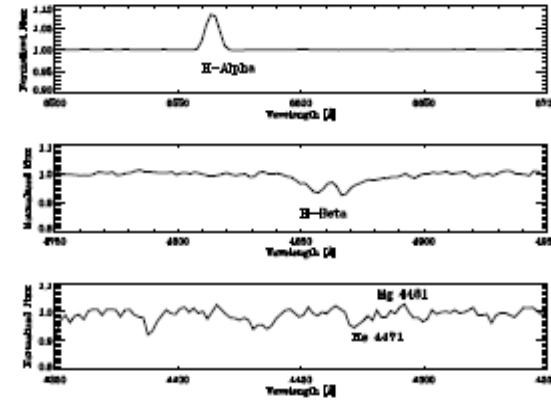
## Stellar seismology

CLEANEST spectrum of CoRoT light-curve



**Fig. 6.** Cleanest Spectra between  $2.4$  and  $2.5$   $c\ d^{-1}$ , from top to bottom: the light curve (lc), lc without orbit perturbations, lc with interpolated gaps, reconstructed time series.

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**Fig. 8.** Close-ups of interesting spectral regions. From top to bottom: H $\alpha$ , H $\beta$ , He 4471  $\text{\AA}$ /Mg 4481  $\text{\AA}$

CoRoT ID 102761769 → →

- Physical parameters:  $V= 13$  B5-6 IV-Ve,  
 $V_{\sin i} = 120$  km/s
  - Frequency spectrum
  - 2 dominant frequencies ↗ NRP multiplet
- (Emilio et al. 2011;  
Andrade et al. 2012)





# SCIENCE

## Stellar rotation (de Medeiros et. al.)

Semi-sinusoidal variability in the light curves

☒☒ rotation periods ?

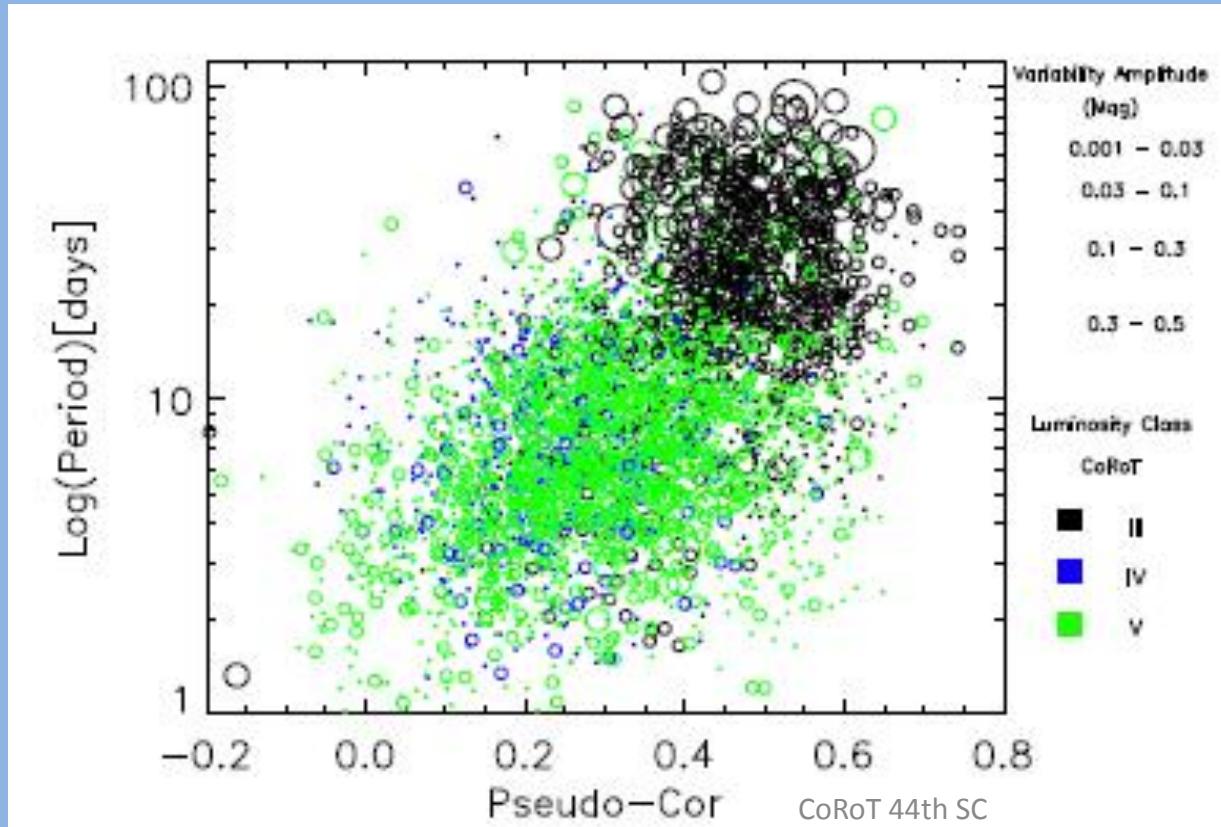
Distribution of stars in  $P \times$  color diagram  
depends strongly on **reddening correction**



# SCIENCE

## Stellar rotation (Medeiros et. al.)

Period - reddened corrected color (size of circle  $\propto$  period)



de Medeiros et al.  
A&A submitted





## SCIENCE

Physics of stellar spots (Silva-Valio et al.)  
(planet eclipses spot during transit)

- ◆ Spot characteristics: (Silva 2003)
  - Size (area coverage); Intensity – temperature; Location (long & lat)
- ◆ Stellar properties from multiple transits
  - Rotation period (Silva-Valio 2008)
  - Differential rotation (Silva-Valio et al. 2010; Silva-Valio & Lanza 2011)
  - Exomoons (Tusnoki & Valio, 2011)
  - Activity cycle



# THANK YOU !